

Substorms in Coupled Models: Expanding Coupling between BATS-R-US and RAM-SCB

D. T. Welling¹

J. Haiducek¹

G. Toth¹

V. K. Jordanova²

¹University of Michigan CLaSP Department

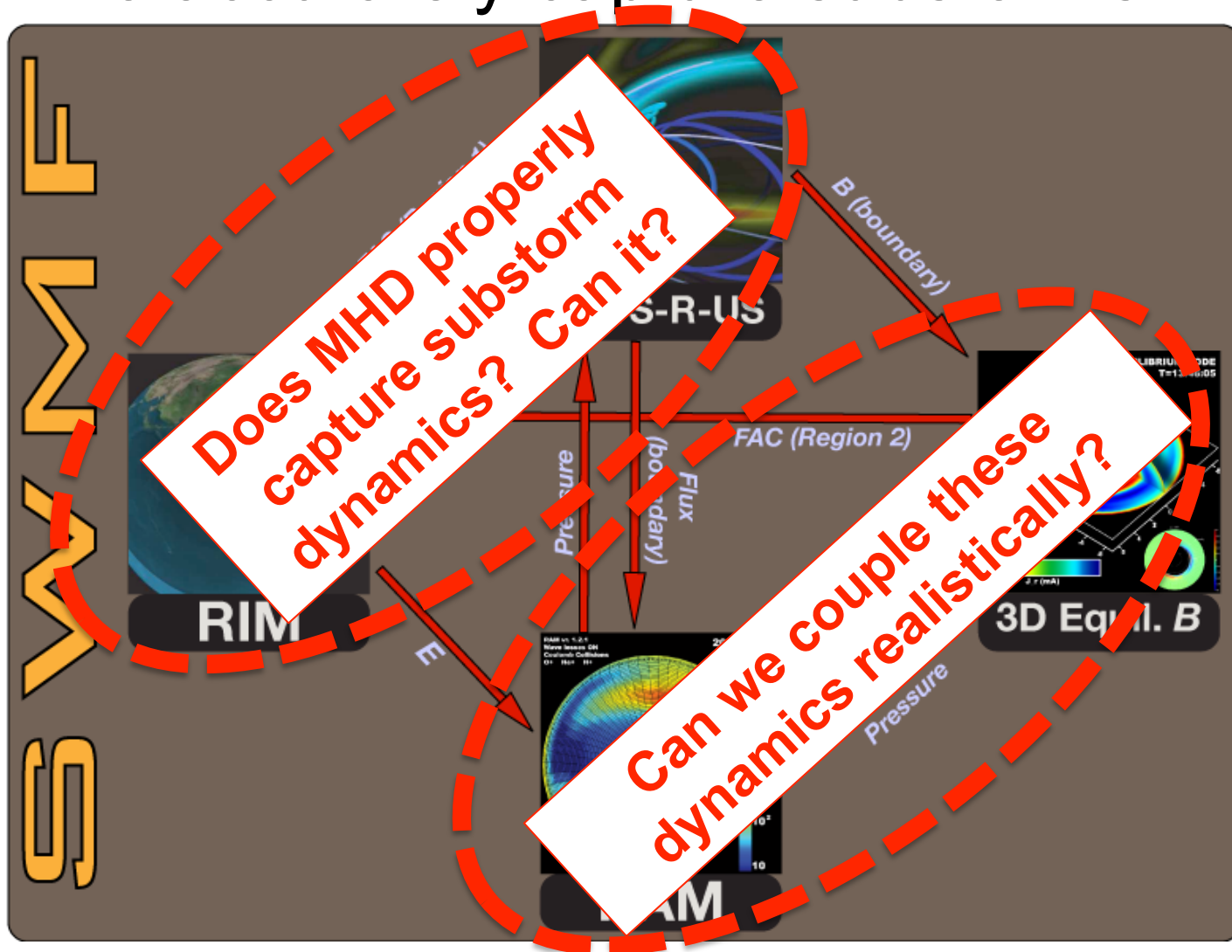
²Los Alamos National Laboratory



Overview



Goal: Used a coupled, multi-scale, global model to accurately capture substorms

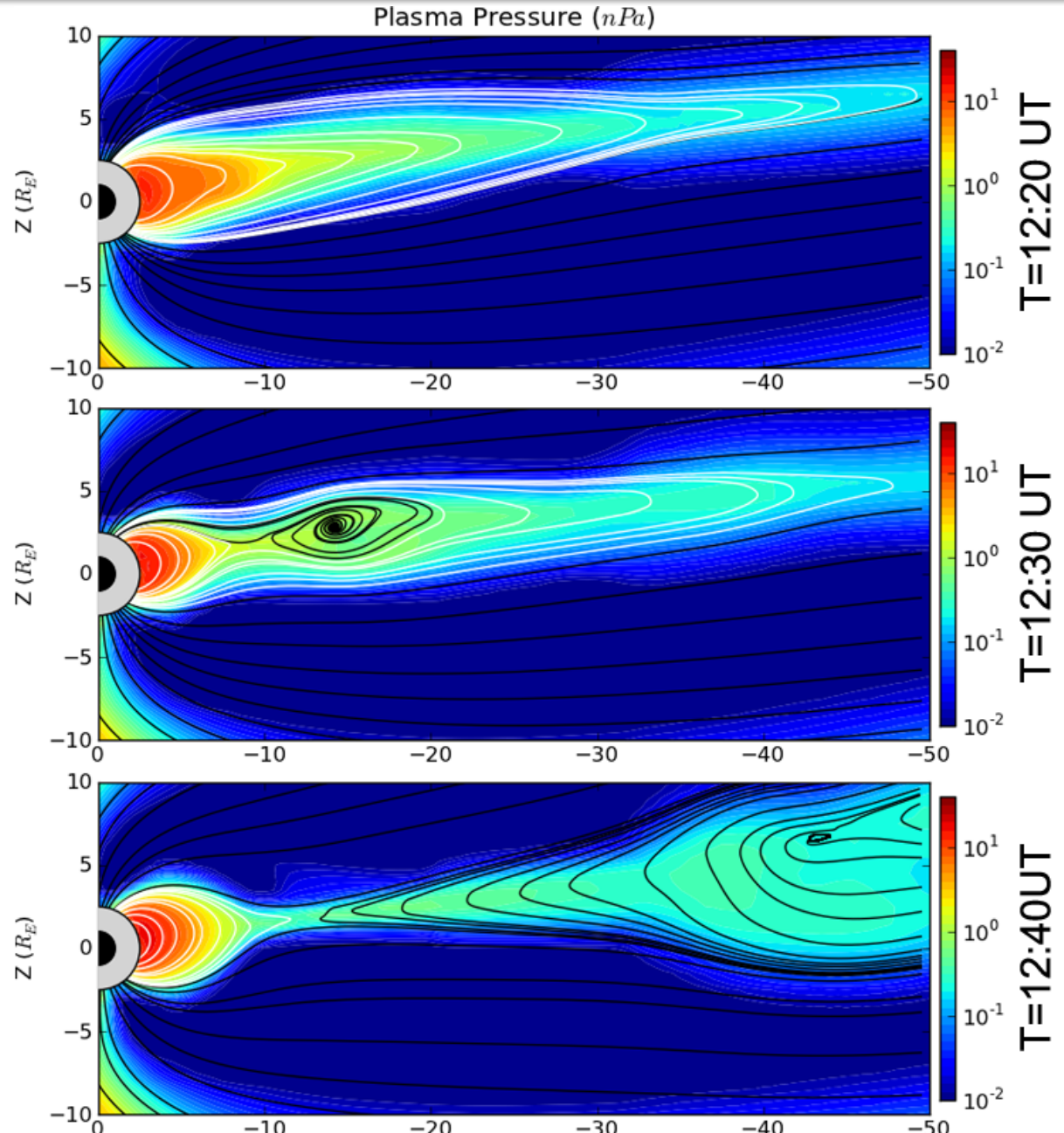


Example “Hit”



Success!

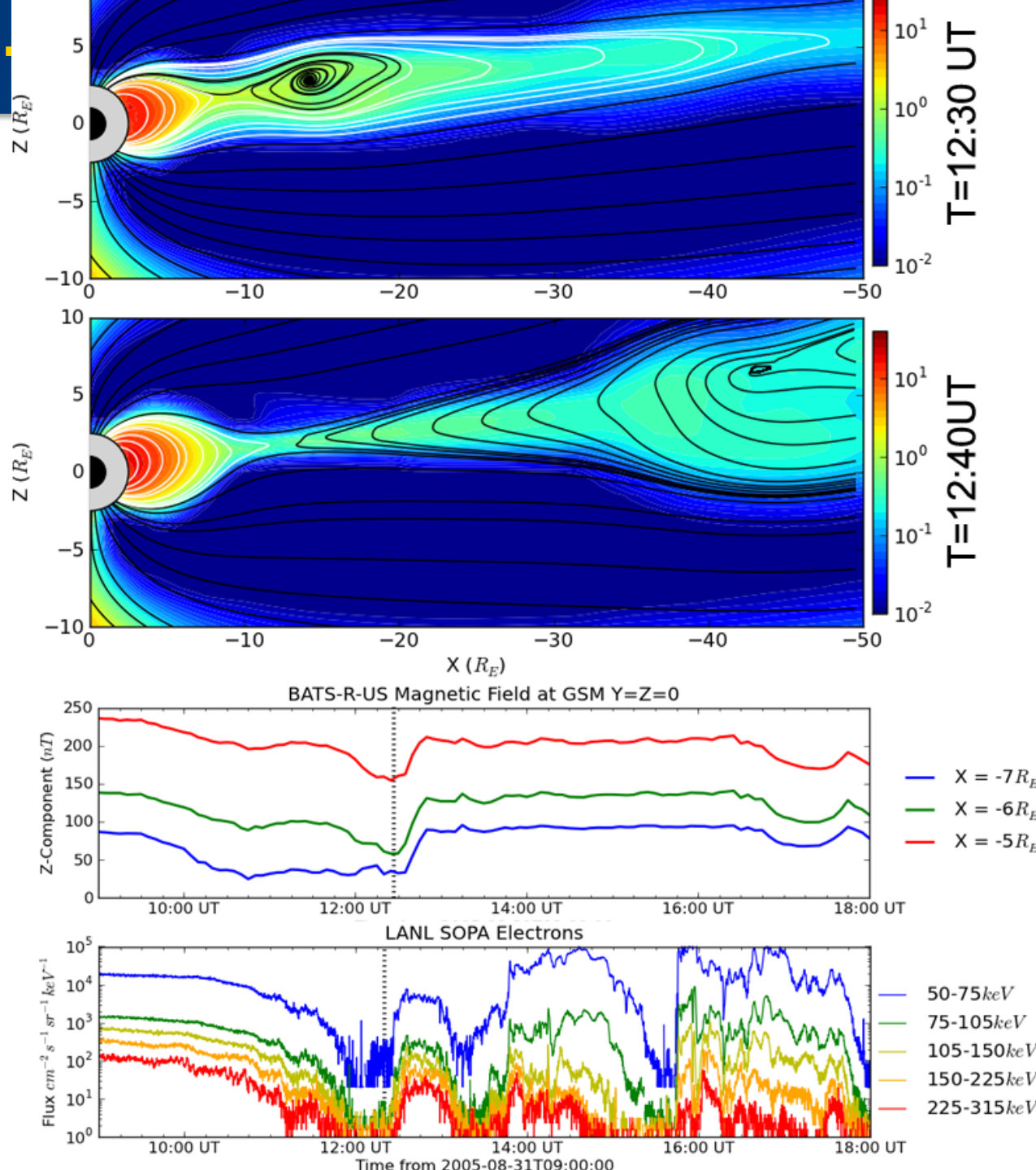
- Southward turning.
- No pressure pulse.
- Good timing agreement with GEO observation



Example “Hit”

Success!

- Southward turning.
- No pressure pulse.
- Good timing agreement with GEO observation

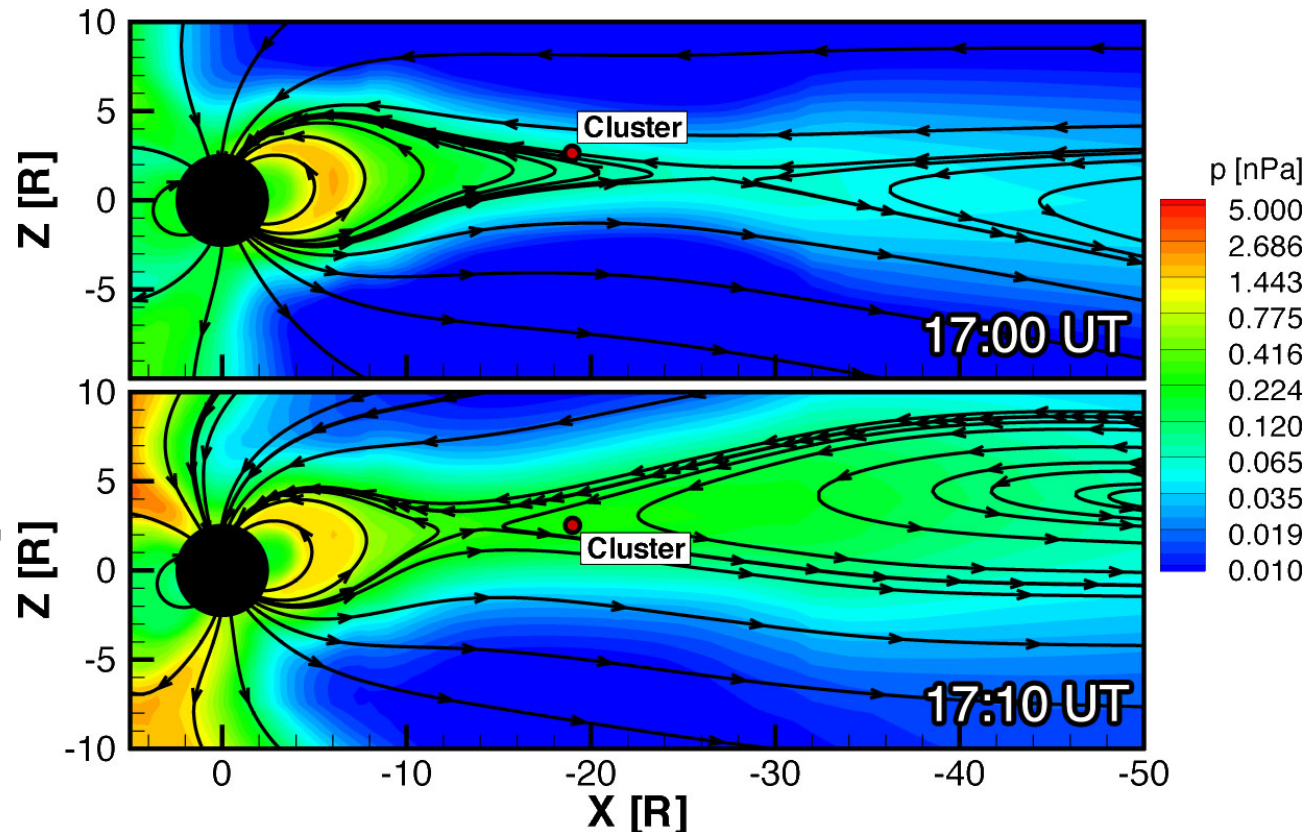
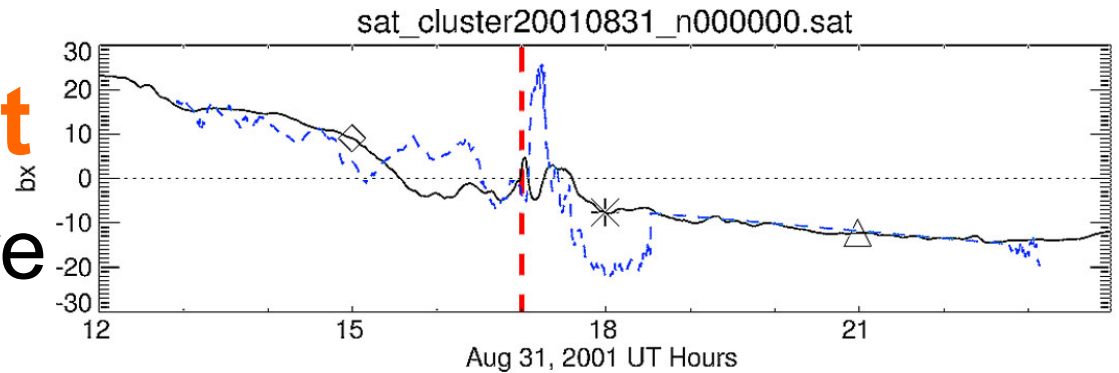


Example “Almost Hit”

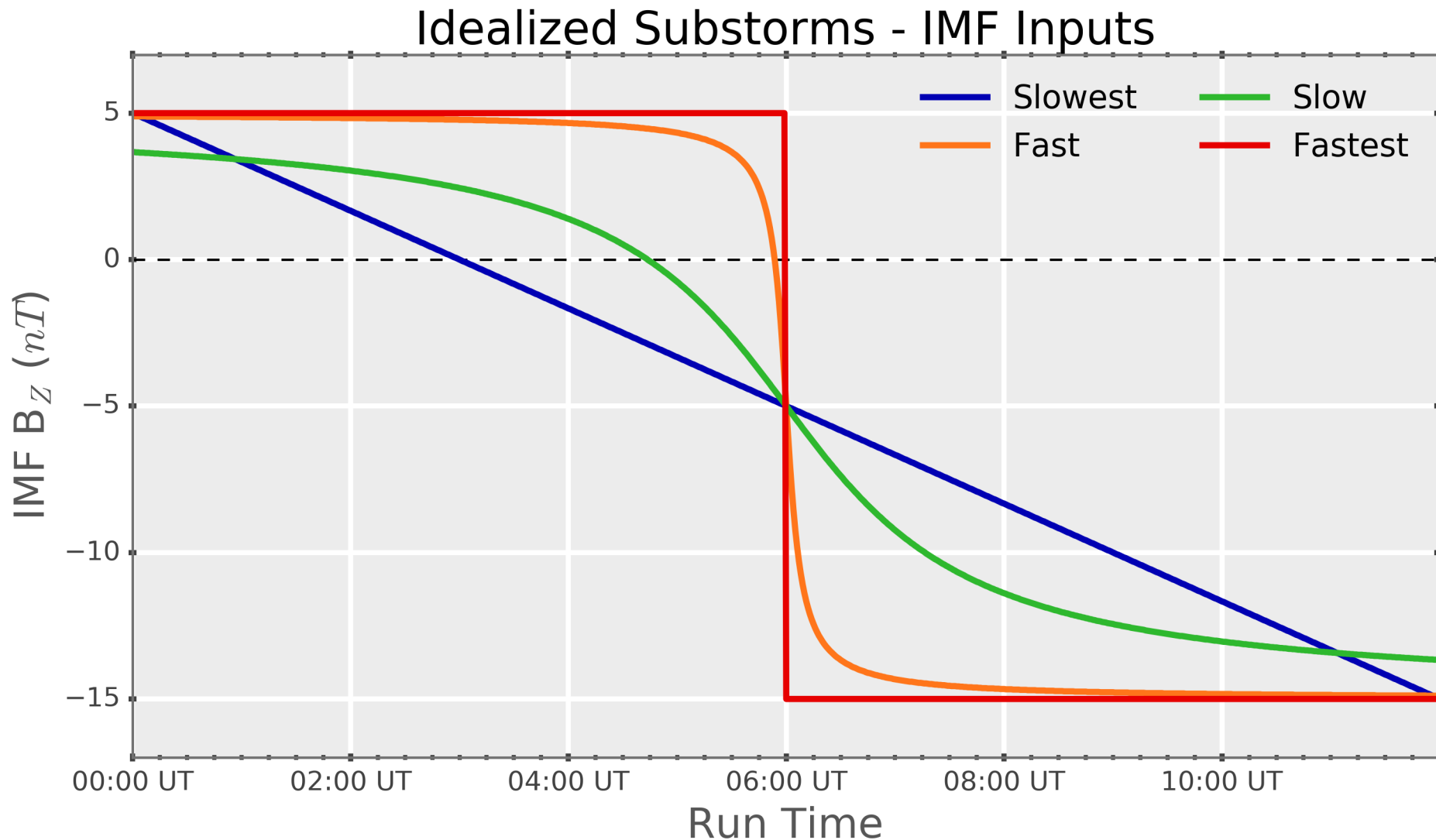


Partial Credit

- Huge pressure pulse.
- Weak, transient southward turning.
- Good timing, poor field geometry.



An Idealized Experiment



Rules for simple loading-unloading:

- Solar wind power input at magnetopause P accumulates energy in magnetotail E .

$$\frac{dE}{dt} = P = \varepsilon(v, B)$$

- Unique minimum energy state for magnetosphere F exists for given solar wind state P .

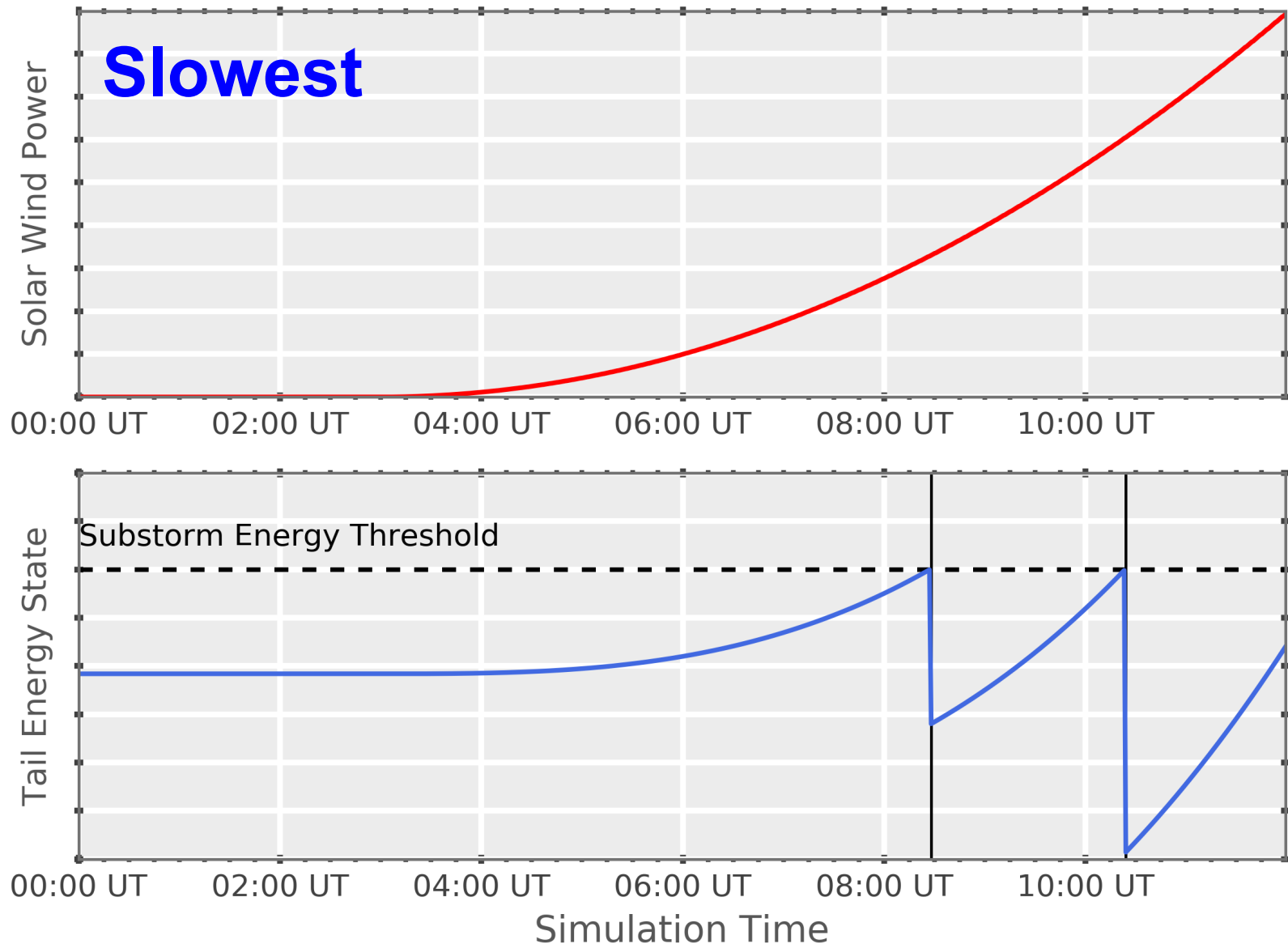
$$F = C - DP$$

- Magnetotail can only move to lower energy state F when energy threshold C is exceeded.

$$E \rightarrow F \quad \text{when} \quad E \geq C$$

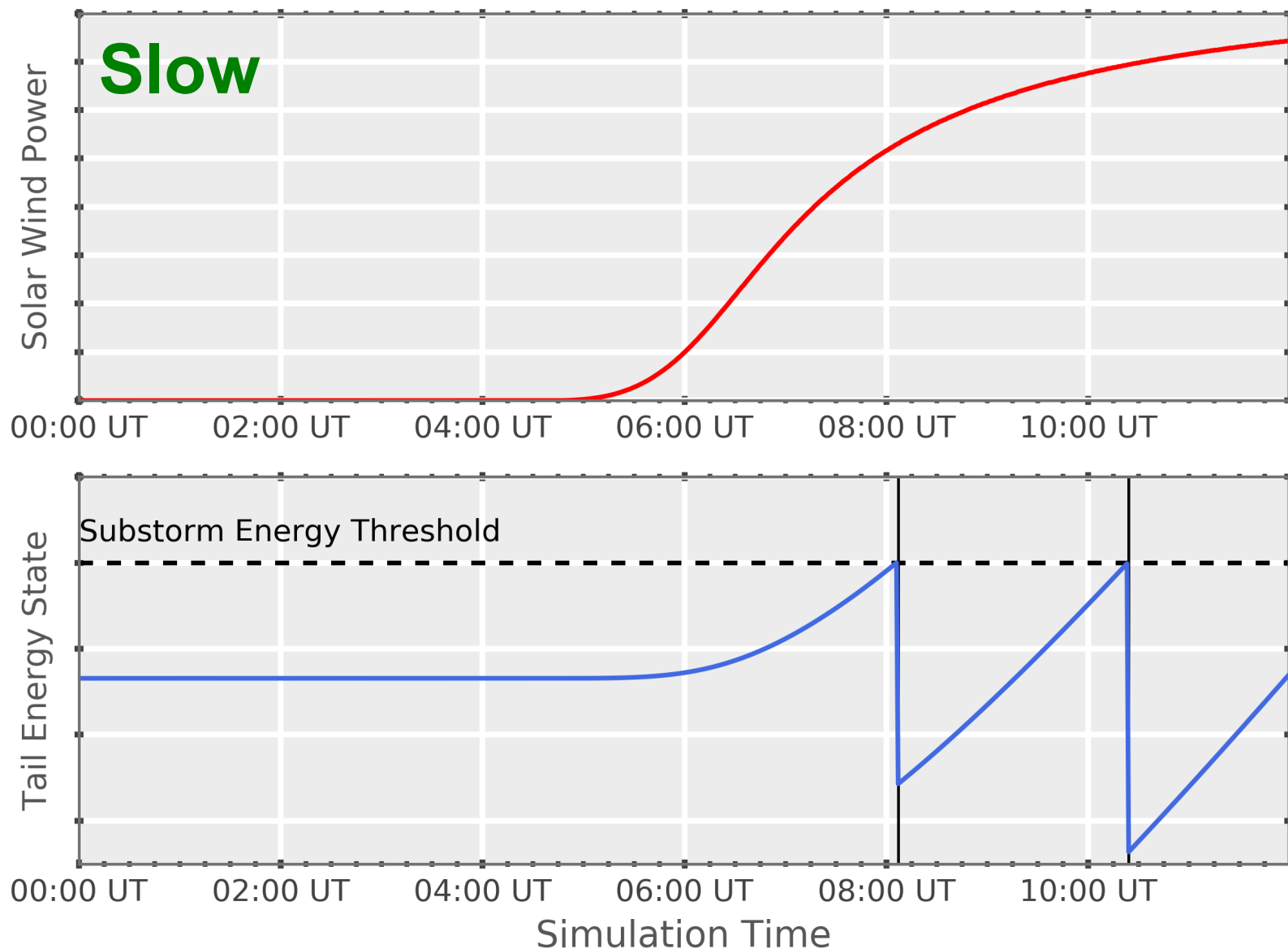
- For constant input, substorms occur with constant period of D
 - Expect $D = 2.5\text{-}3$ hr [Borovsky et al., JGR, 1993; Freeman and Farrugia, JGR, 1995; Huang et al., JGR, 2003]

Expected Results



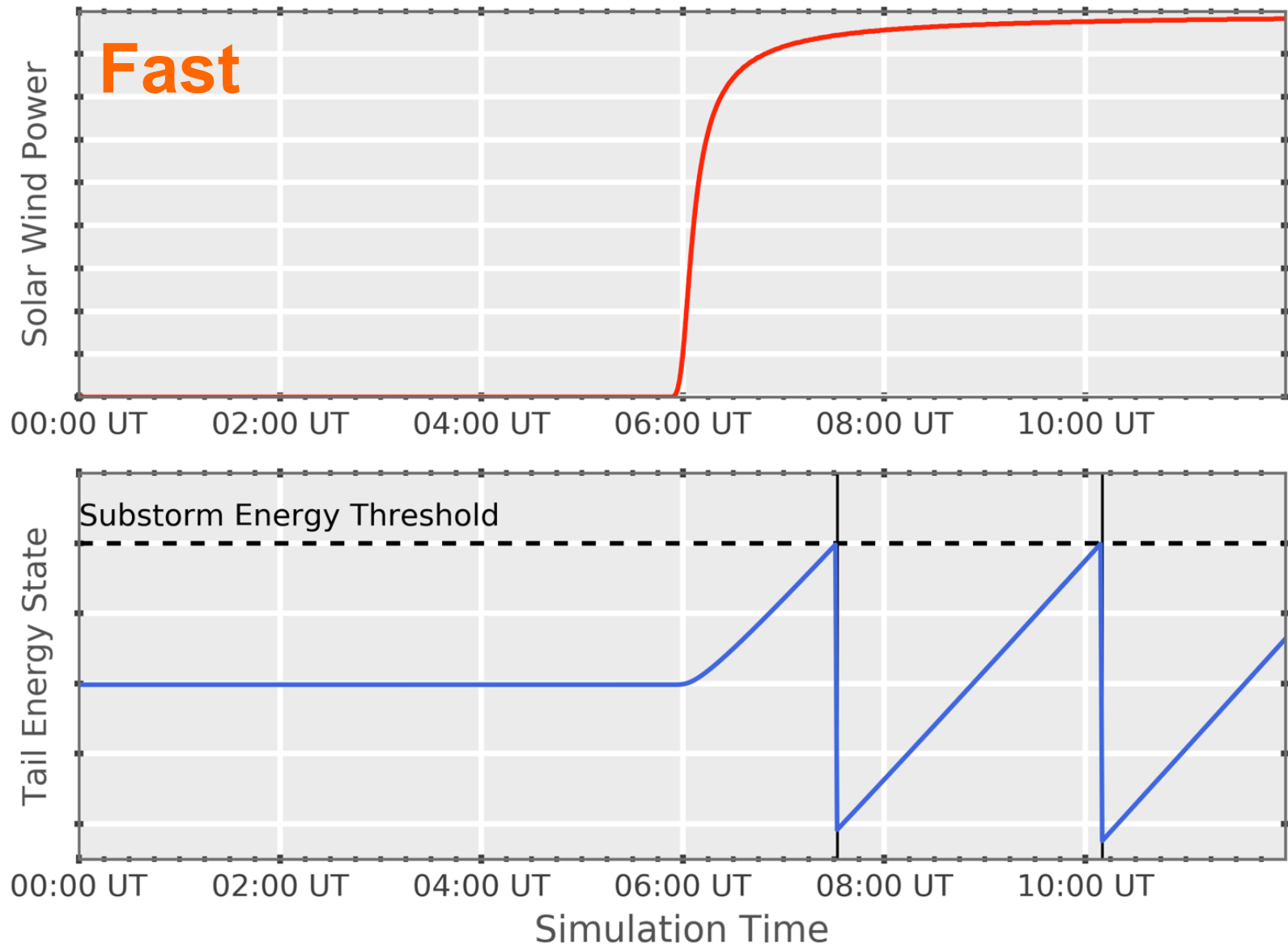
Freeman, M.P. & Morley, S.K., 2004. A minimal substorm model that explains the observed statistical distribution of times between substorms. *Geophysical Research Letters*, 31(12), p.L12807.

Expected Results



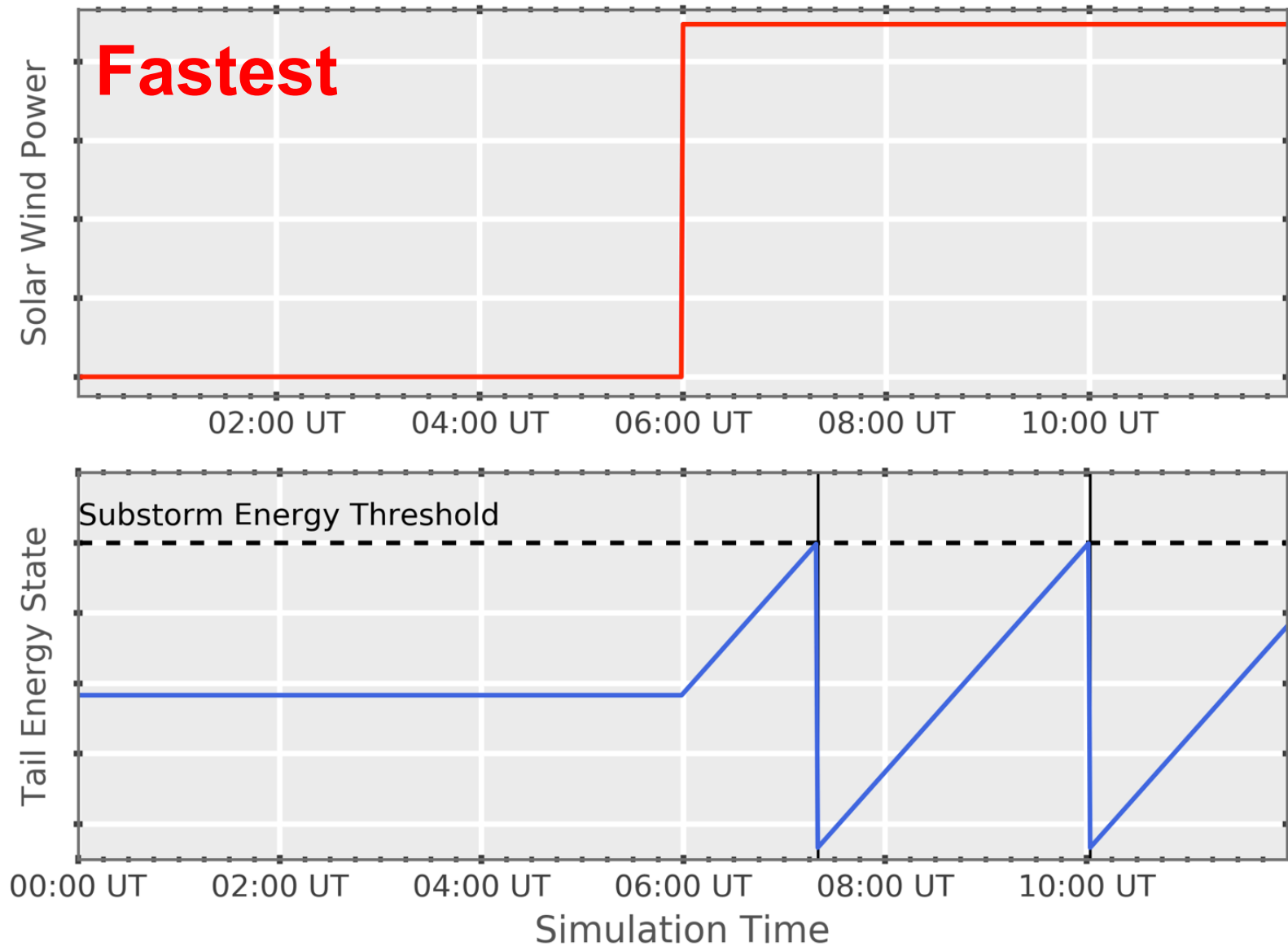
Freeman, M.P. & Morley, S.K., 2004. A minimal substorm model that explains the observed statistical distribution of times between substorms. *Geophysical Research Letters*, 31(12), p.L12807.

Expected Results



Freeman, M.P. & Morley, S.K., 2004. A minimal substorm model that explains the observed statistical distribution of times between substorms. *Geophysical Research Letters*, 31(12), p.L12807.

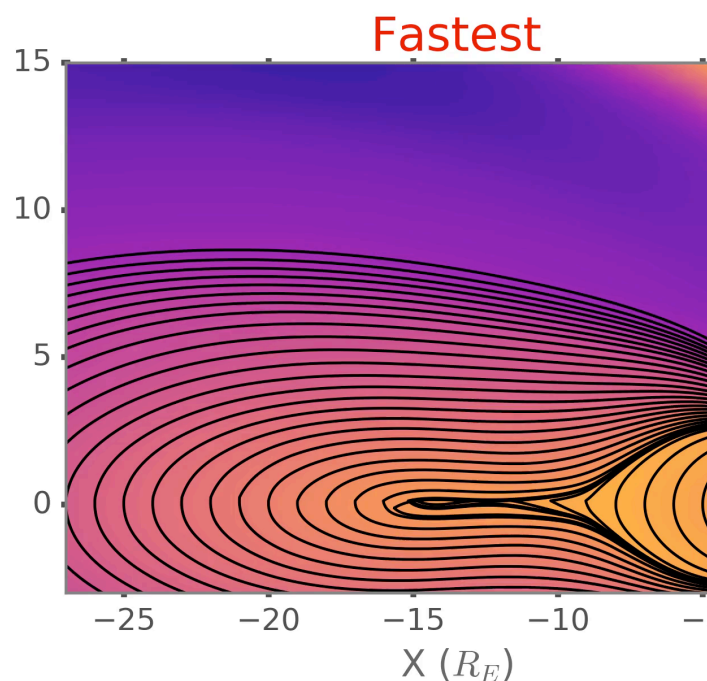
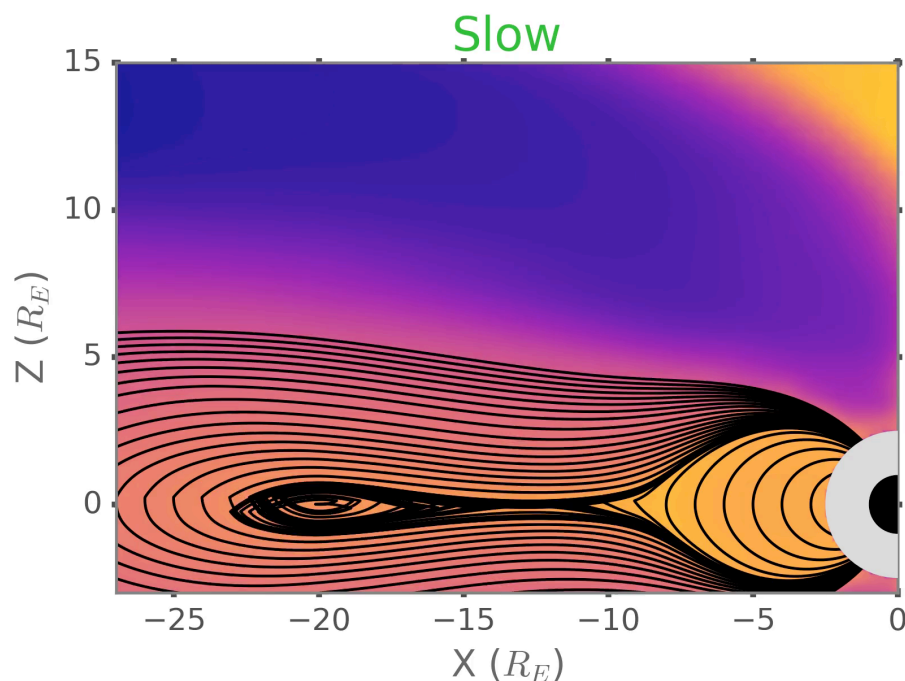
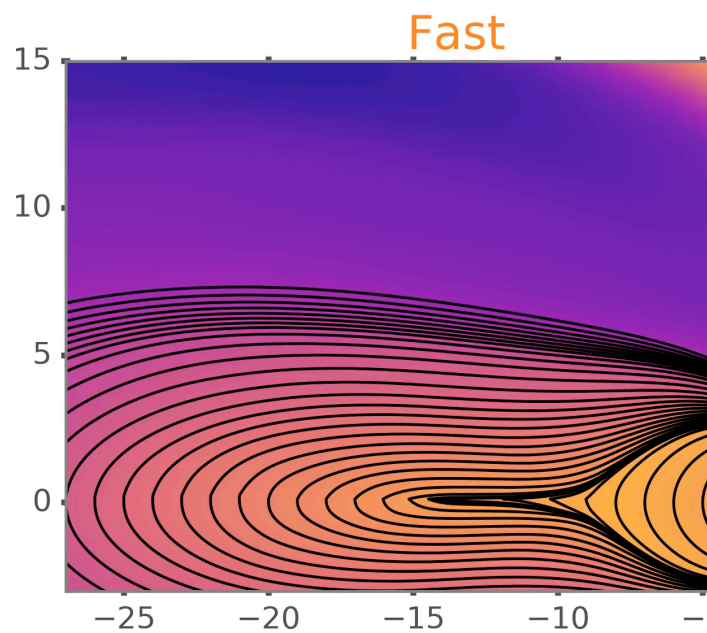
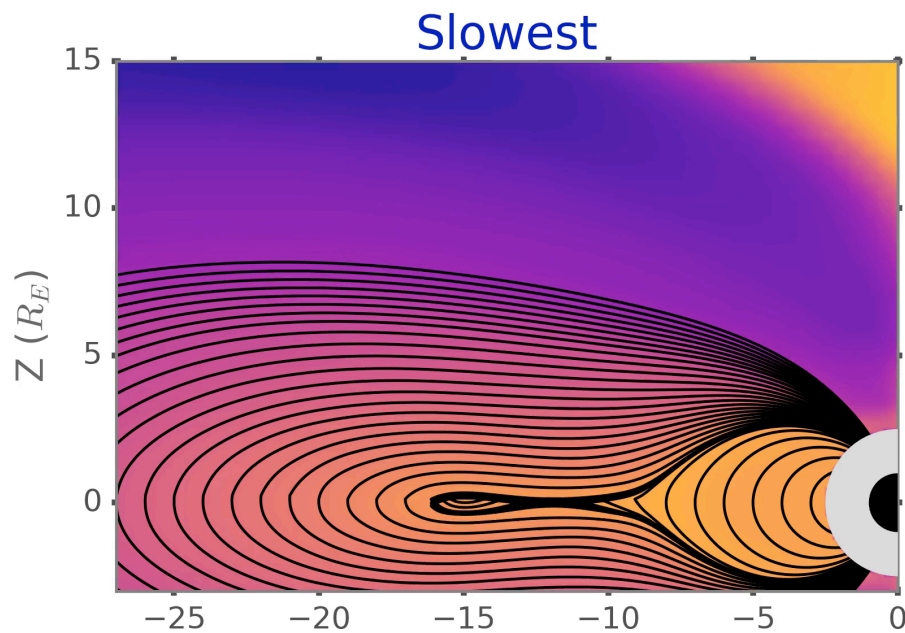
Expected Results



Freeman, M.P. & Morley, S.K., 2004. A minimal substorm model that explains the observed statistical distribution of times between substorms. *Geophysical Research Letters*, 31(12), p.L12807.

Ideal

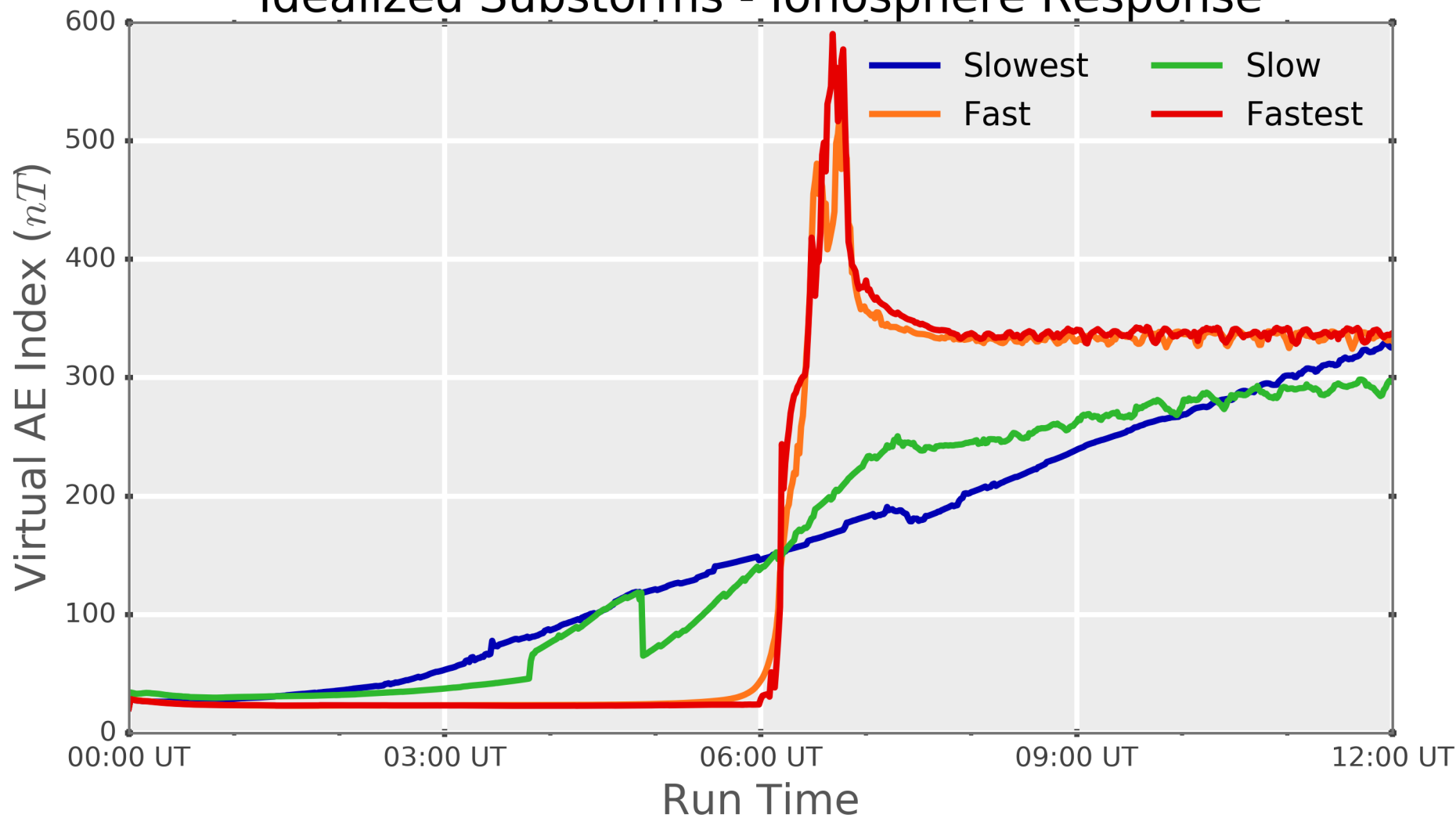
Run Time = 01h00m



Results - Ionosphere



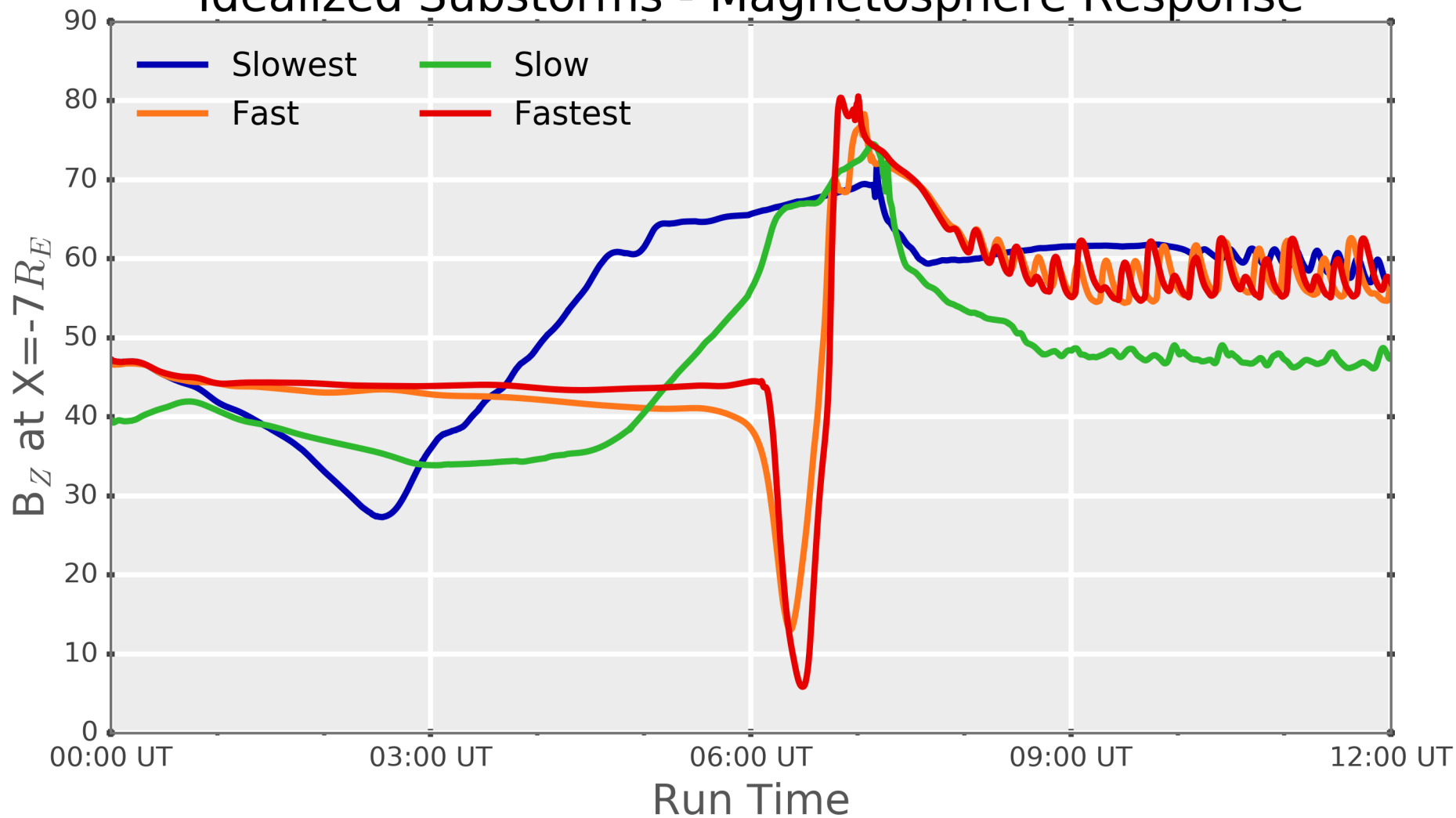
Idealized Substorms - Ionosphere Response



Results - Magnetosphere



Idealized Substorms - Magnetosphere Response



Under what conditions can ideal MHD capture substorms?

- System appears to be “leaky”.
- Quick timescale IMF change required.
- What about real world conditions?

How can MHD be extended to improve performance during substorms?

- Resistivity: Anomalous, Hall
- Heavy Ions: multi species/fluid
- Anisotropic MHD
- MHD with embedded Particle-in-cell (EPIC)

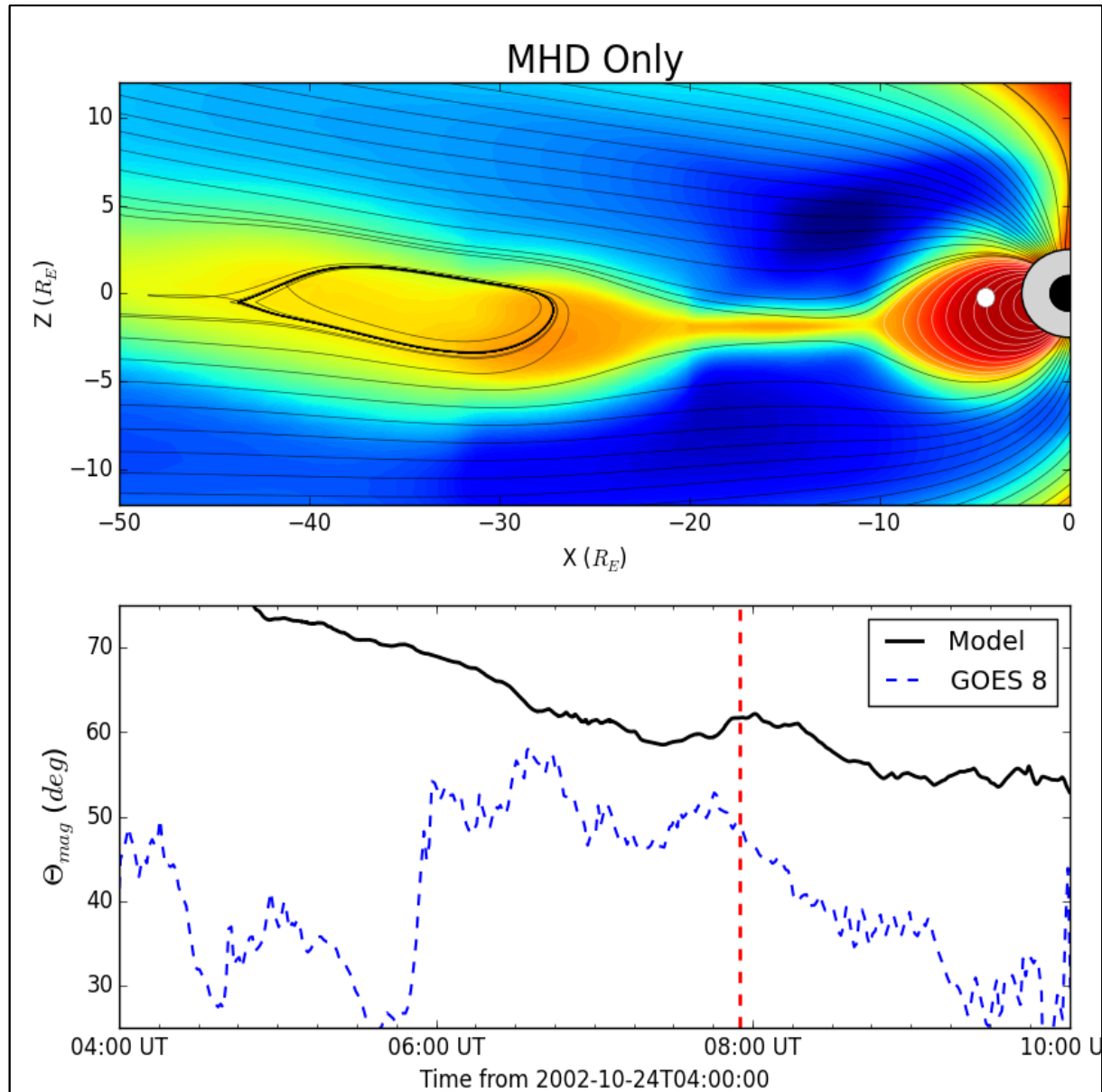
Outflow & Substorms



MHD Only:
weak event,
terrible timing.

MHD+PWOM:
Stronger
event, poor
timing.

MHD+GPW:
Excellent
comparison.



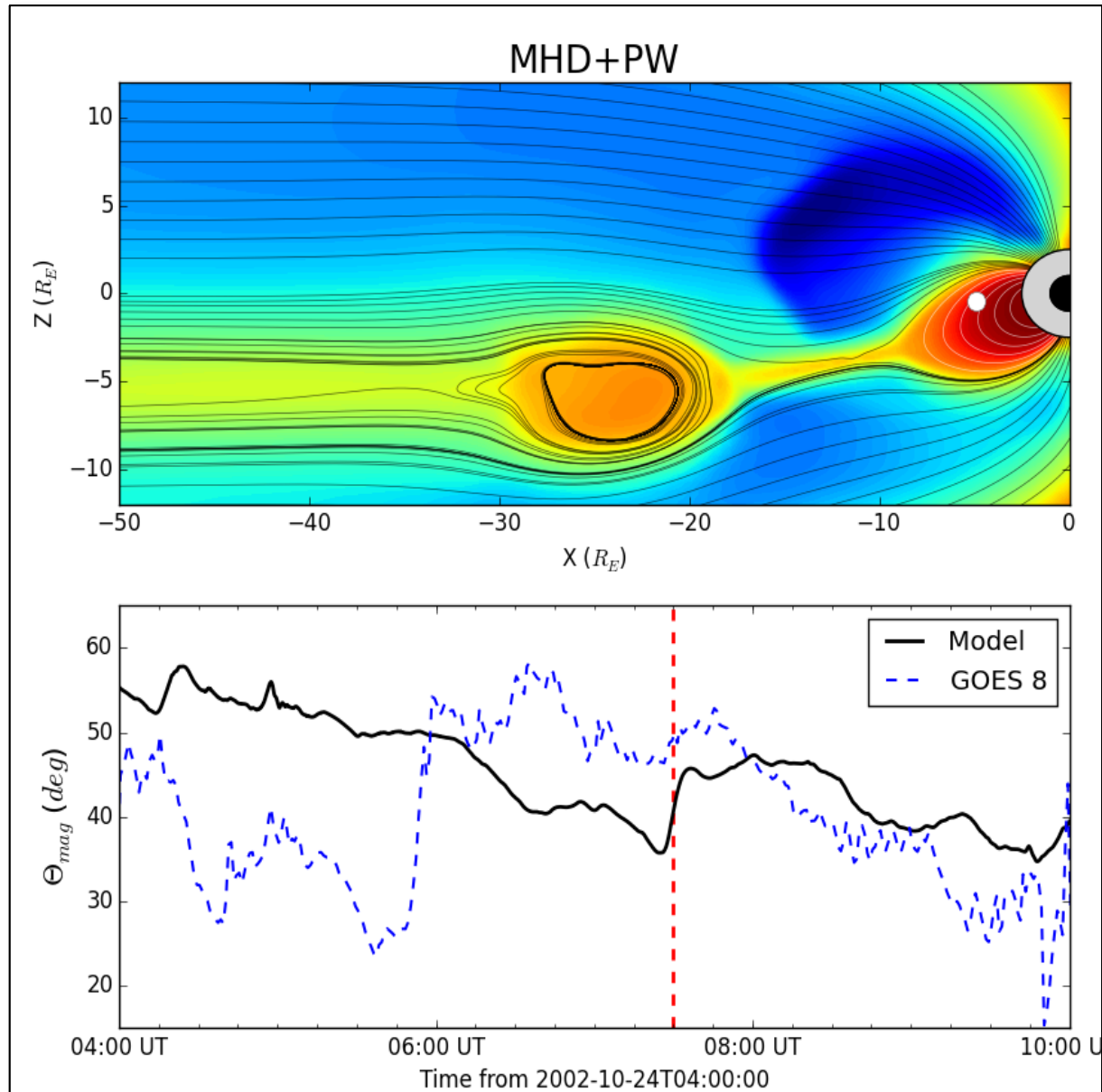
Outflow & Substorms



MHD Only:
weak event,
terrible timing.

MHD+PWOM:
Stronger
event, poor
timing.

MHD+GPW:
Excellent
comparison.



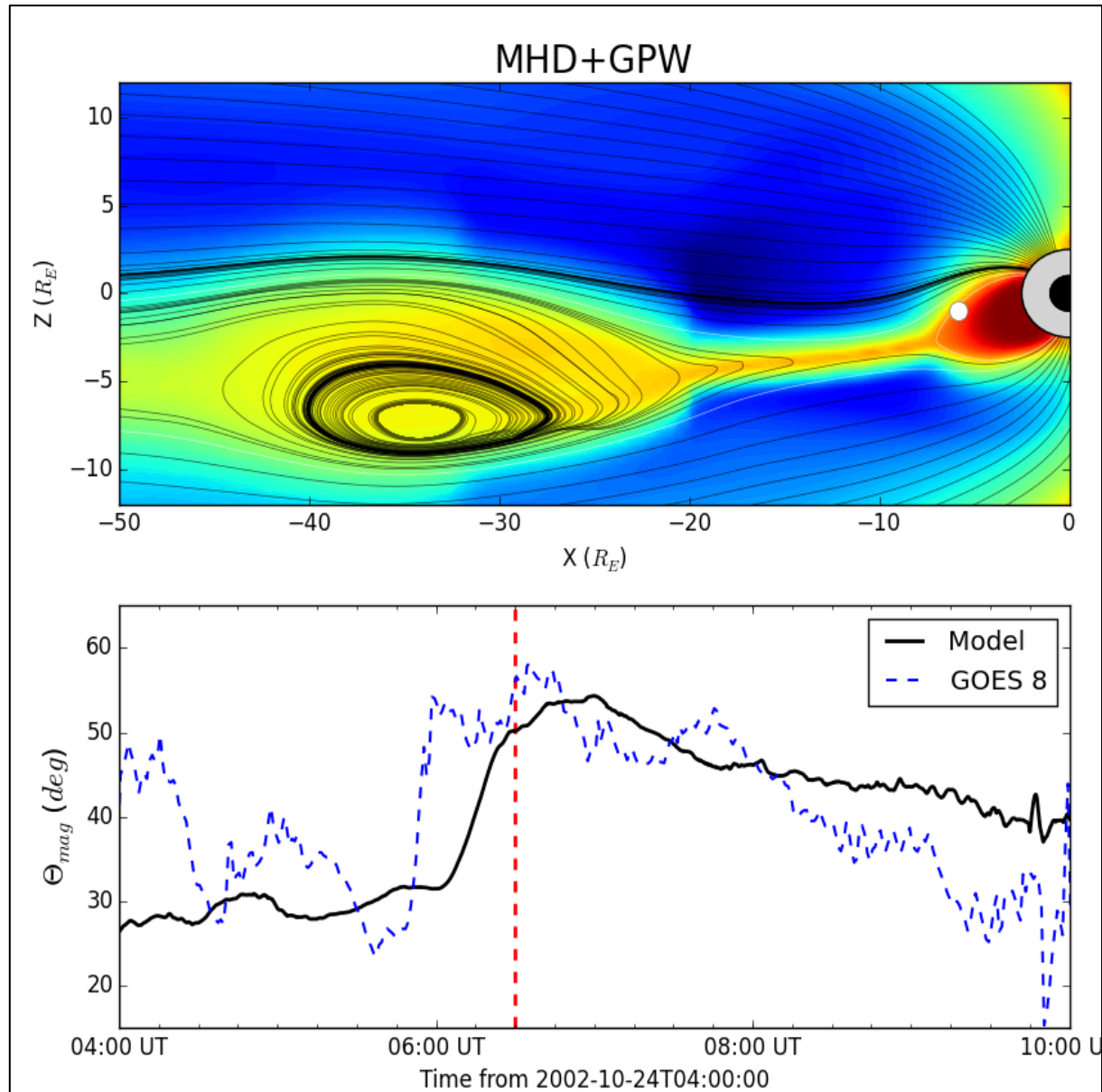
Outflow & Substorms



MHD Only:
weak event,
terrible timing.

MHD+PWOM:
Stronger
event, poor
timing.

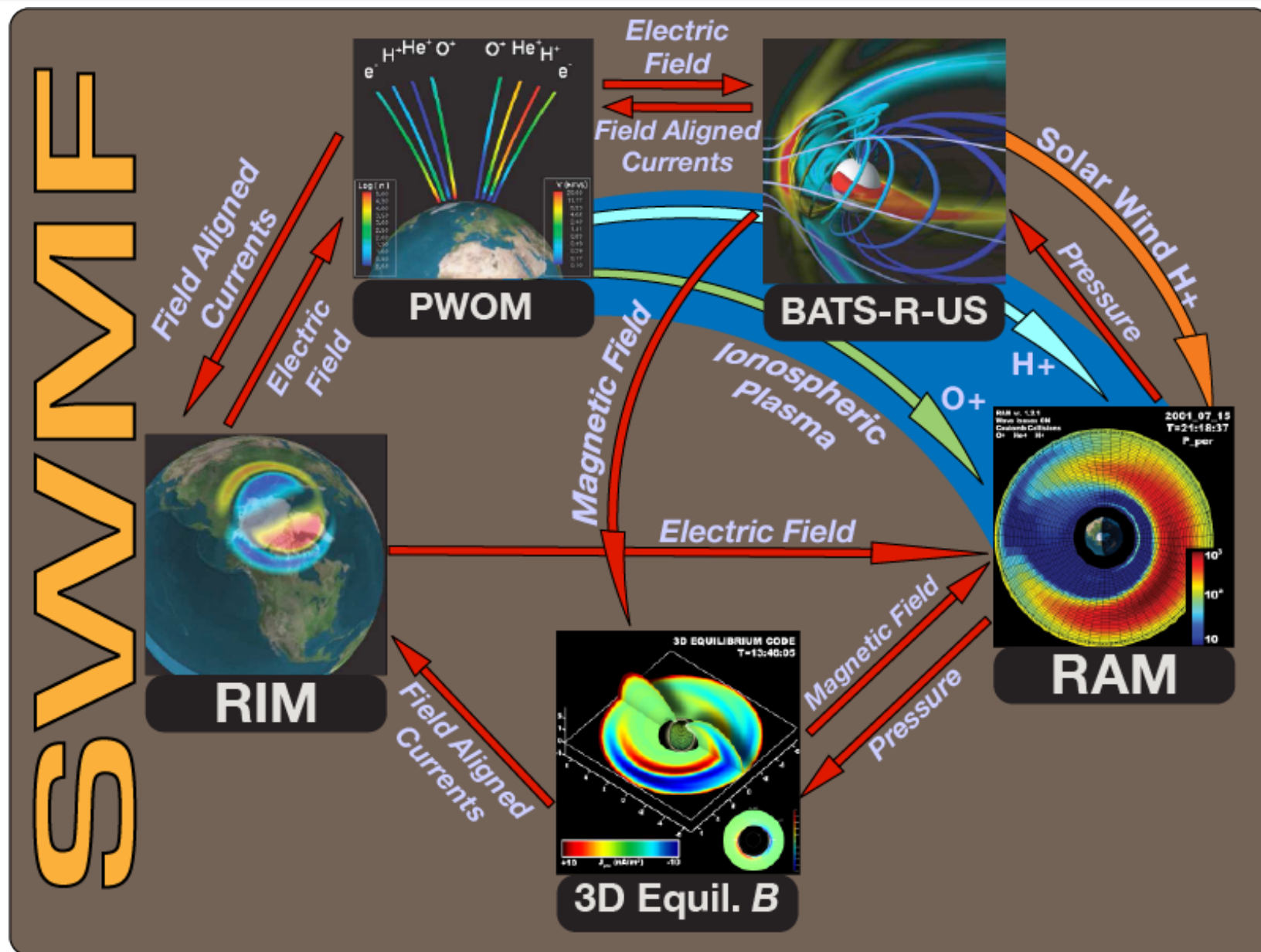
MHD+GPW:
Excellent
comparison.



SWMF: GM-IM-IE w/ RAMSCB



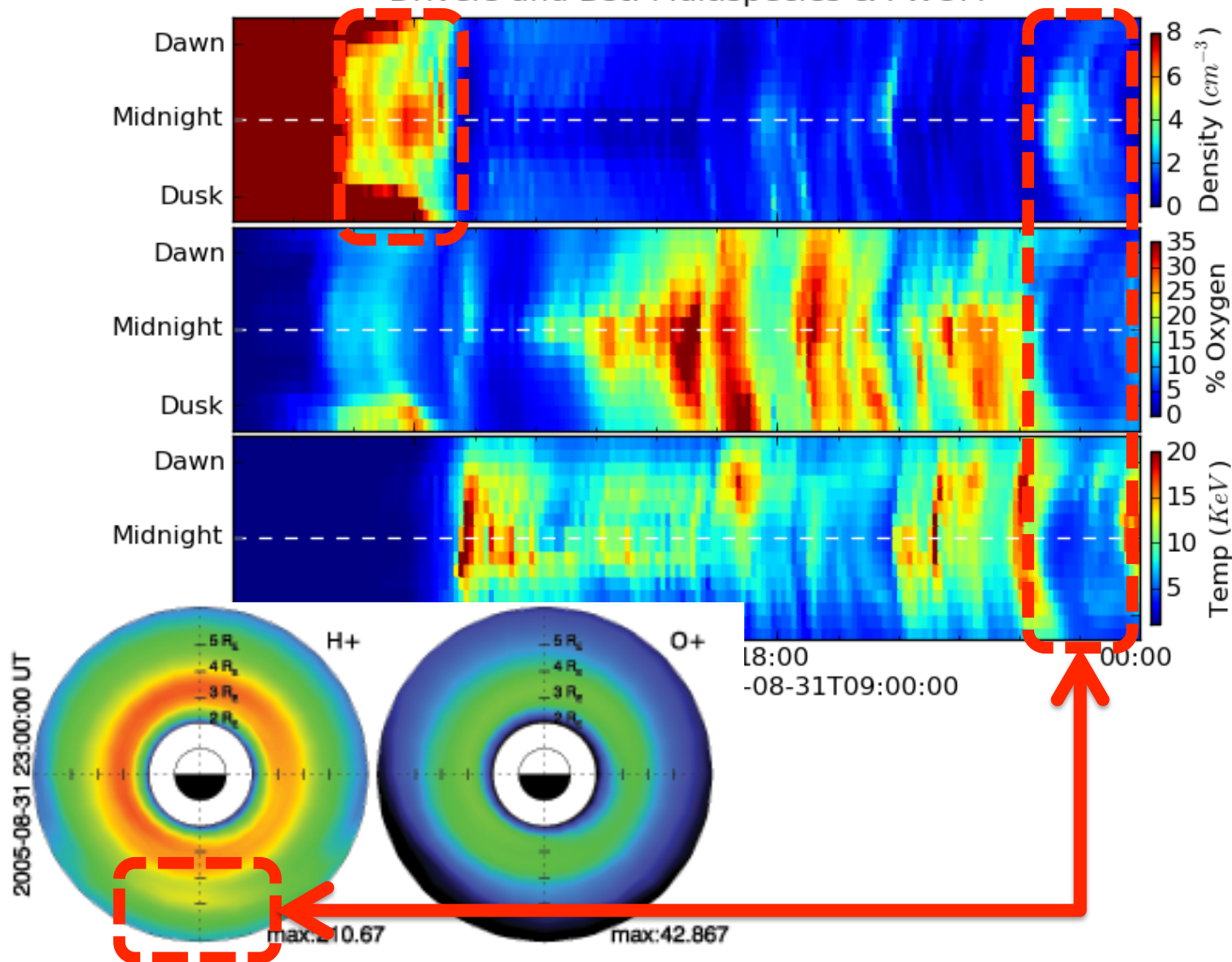
Expanded Earth Coupling



Substorm Injections



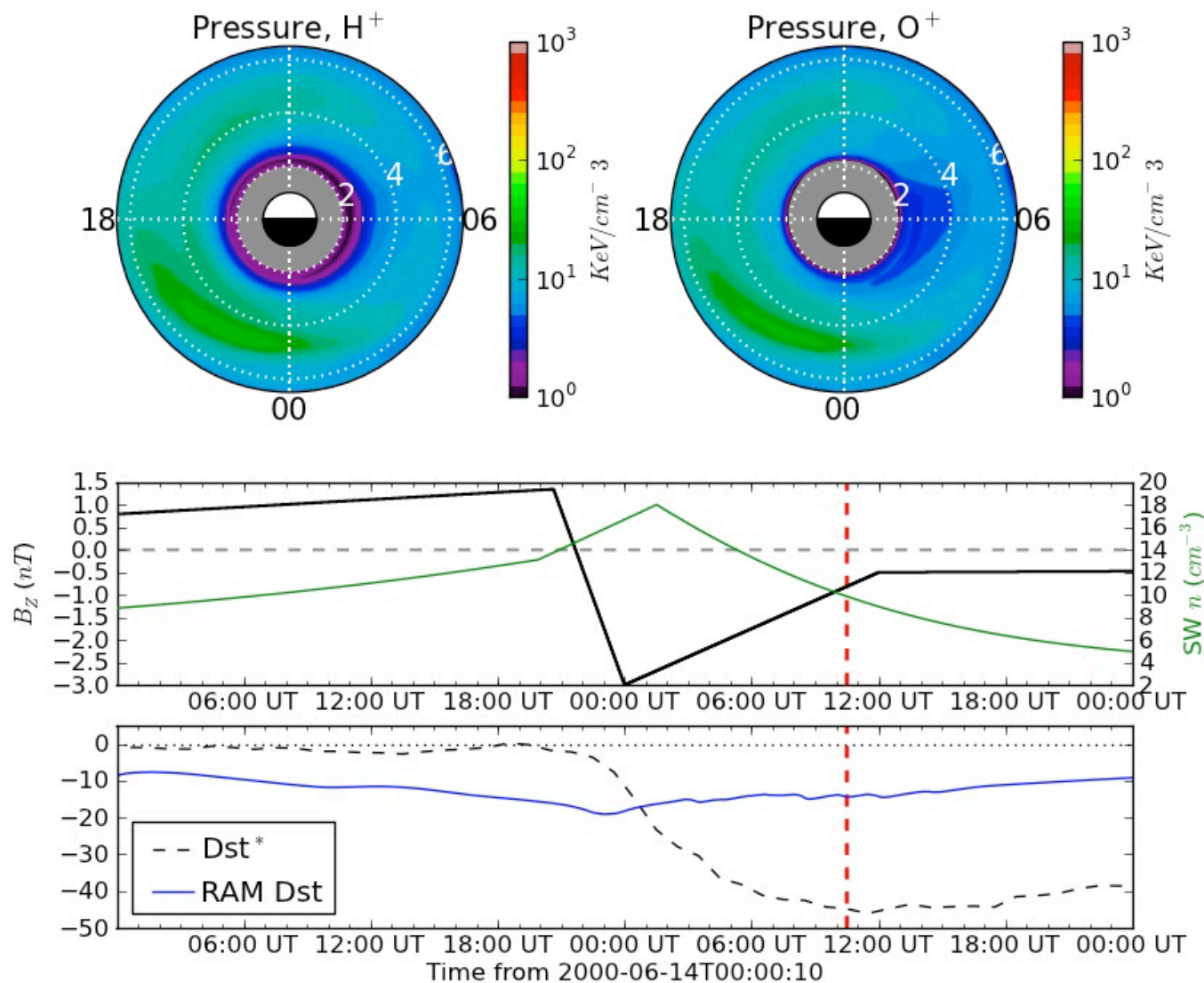
Drivers and Dst: Multispecies & PWOM



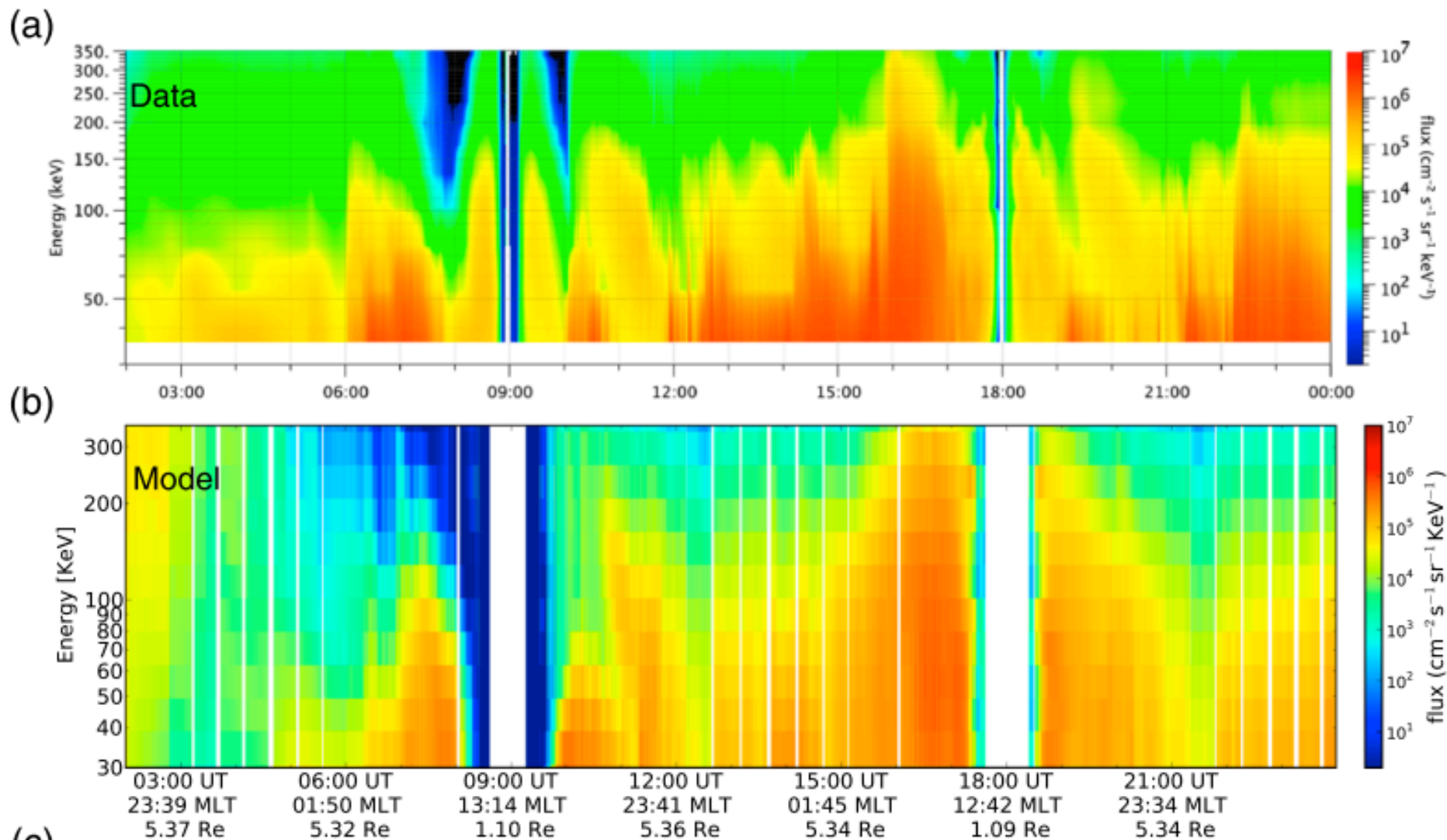
Ring Current Injections



SEA Time = +10:30



Injections into RAM-SCB

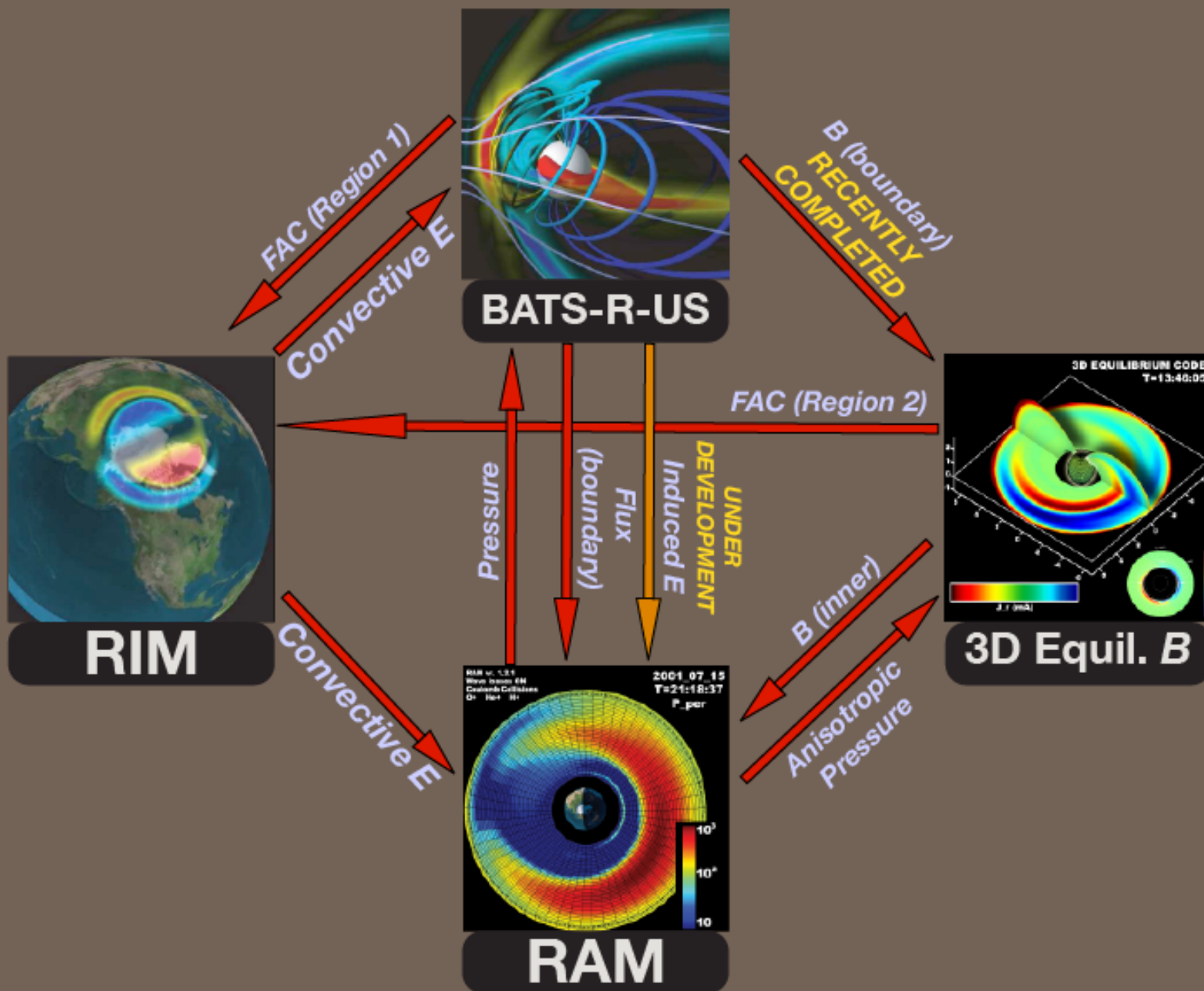


Yu, Y. et al., 2014. The role of ring current particle injections: Global simulations and Van Allen Probes observations during 17 March 2013 storm. *Geophysical Research Letters*, 41(4), pp.1126–1132.

Induced E-Field Coupling



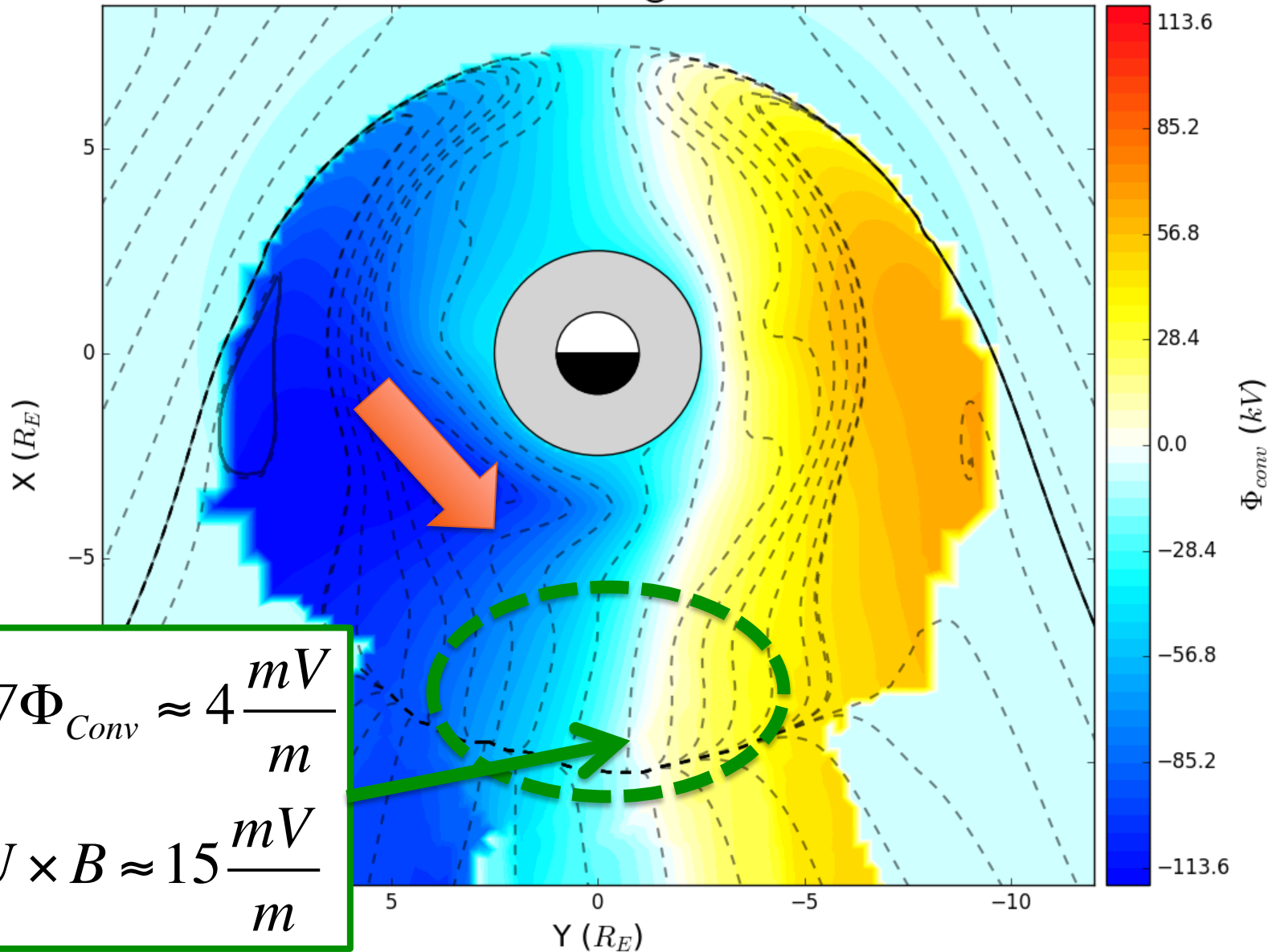
Earth
Magnetosphere
Coupling
System



Potential Field vs. Velocity



Convection Potential @ T=06h40m00



$$\nabla \Phi_{Conv} \approx 4 \frac{mV}{m}$$

$$U \times B \approx 15 \frac{mV}{m}$$

Ideal MHD can reproduce substorm features

- Quality and reliability still under evaluation.
- Many option to expand MHD and improve MHD performance during substorm events.

Coupling to ring current models allows substorm dynamics to be driven into the inner magnetosphere

- Particle injections, convection field produce qualitatively reasonable results.
- We are expanding coupling to use the whole $U \times B$ electric field.